

Assuming the Sun to be a spherical body of radius  $R$  at a temperature of  $TK$ , evaluate the total radiant power incident on Earth at a distance  $r$  from the Sun

Where  $r_0$  is the radius of the Earth and  $\sigma$  is Stefan's constant. (JEE MAIN 2006)

**A**  $4\pi r_0^2 R^2 \sigma \frac{T^4}{r^2}$

**B**  $\pi r_0^2 R^2 \sigma \frac{T^4}{r^2}$

**C**  $r_0^2 R^2 \sigma \frac{T^4}{4\pi r^2}$

**D**  $R^2 \sigma \frac{T^4}{r^2}$

$$\begin{aligned} \text{Total power radiated by sun} &= \sigma A T^4 \\ &= \sigma (4\pi R^2) T^4 \end{aligned}$$

$$\begin{aligned} \text{Power intensity at earth's surface} &= \frac{\text{Total Power}}{\text{Total Area}} \\ &= \frac{\sigma 4\pi R^2 T^4}{4\pi r^2} \end{aligned}$$

$$\begin{aligned} \therefore \text{Total power received by earth} &= \text{Power intensity} \times \text{Area} \\ &= \left( \frac{\sigma 4\pi R^2 T^4}{4\pi r^2} \right) (\pi r_0^2) \\ &= \frac{\pi r_0^2 R^2 \sigma T^4}{r^2} \end{aligned}$$